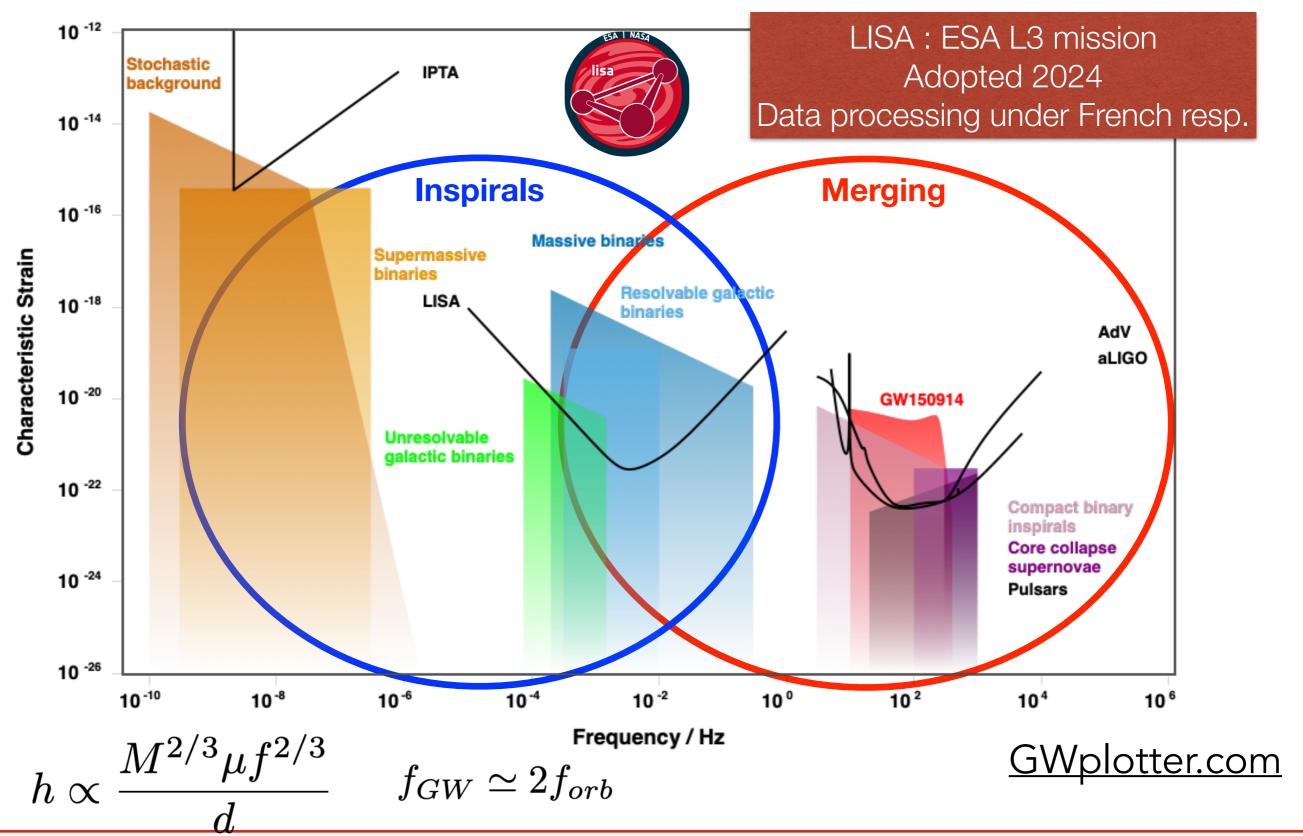


DIFFERENT SOURCES OF GW







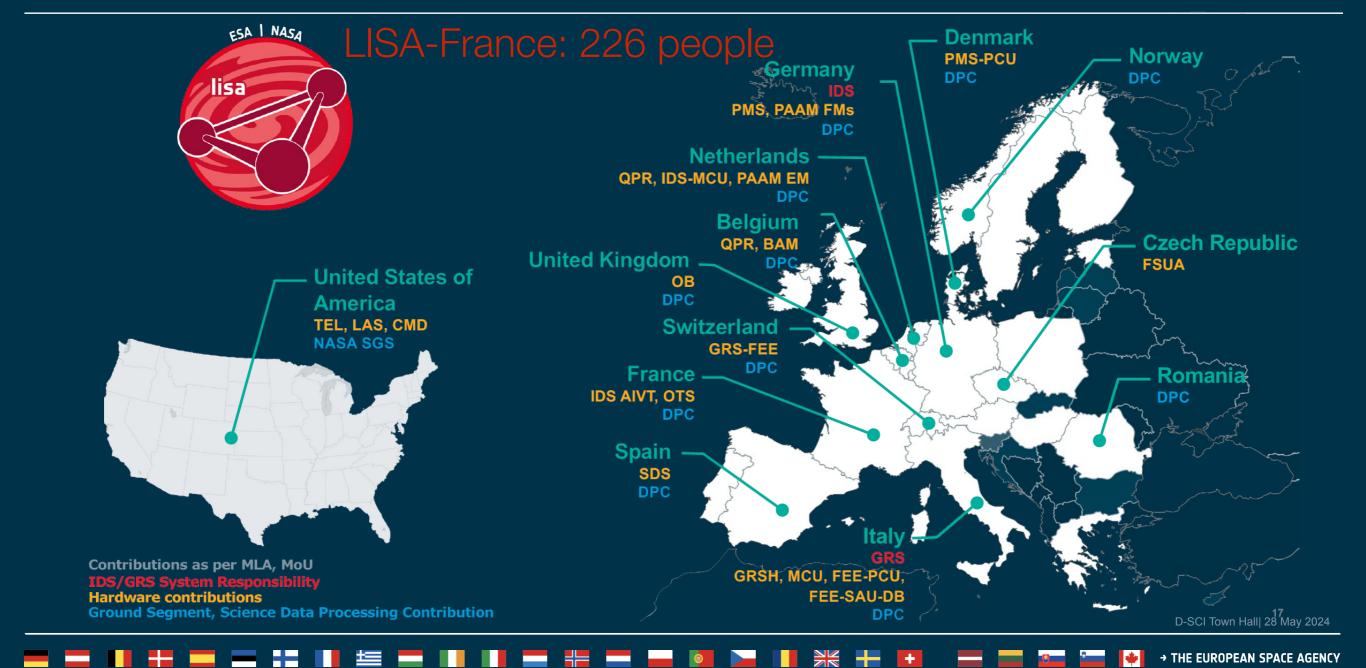


cnrs

LISA COLLABORATION

LISA - An international mission led by ESA



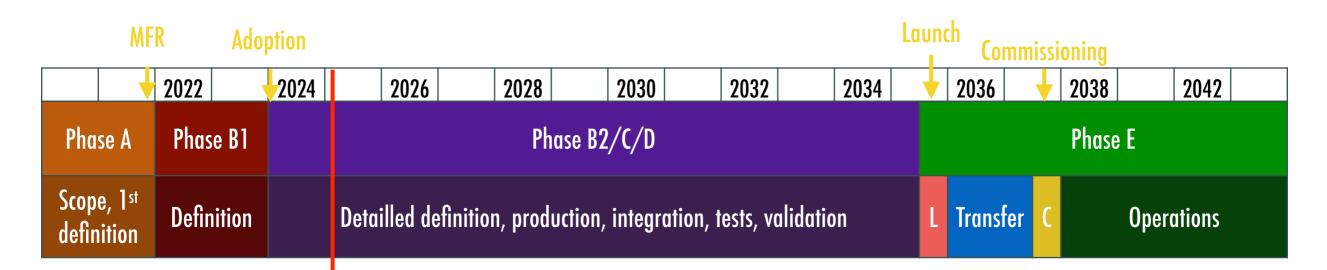






cnrs

TIMELINE AND STATUS



- 1993: first proposal ESA/NASA
- Selection 2017
- End 2021: success of the ESA Mission Formulation Review
- 25/01/2024: adoption
- (New) LISA Science Team in place -> group on content of Data releases
- Long building phase of multiple MOSAs: 6 flight models + test models
- Launch 2035
- 1.5 years of transfer, 4.5 years nominal mission, 6.5 years extension



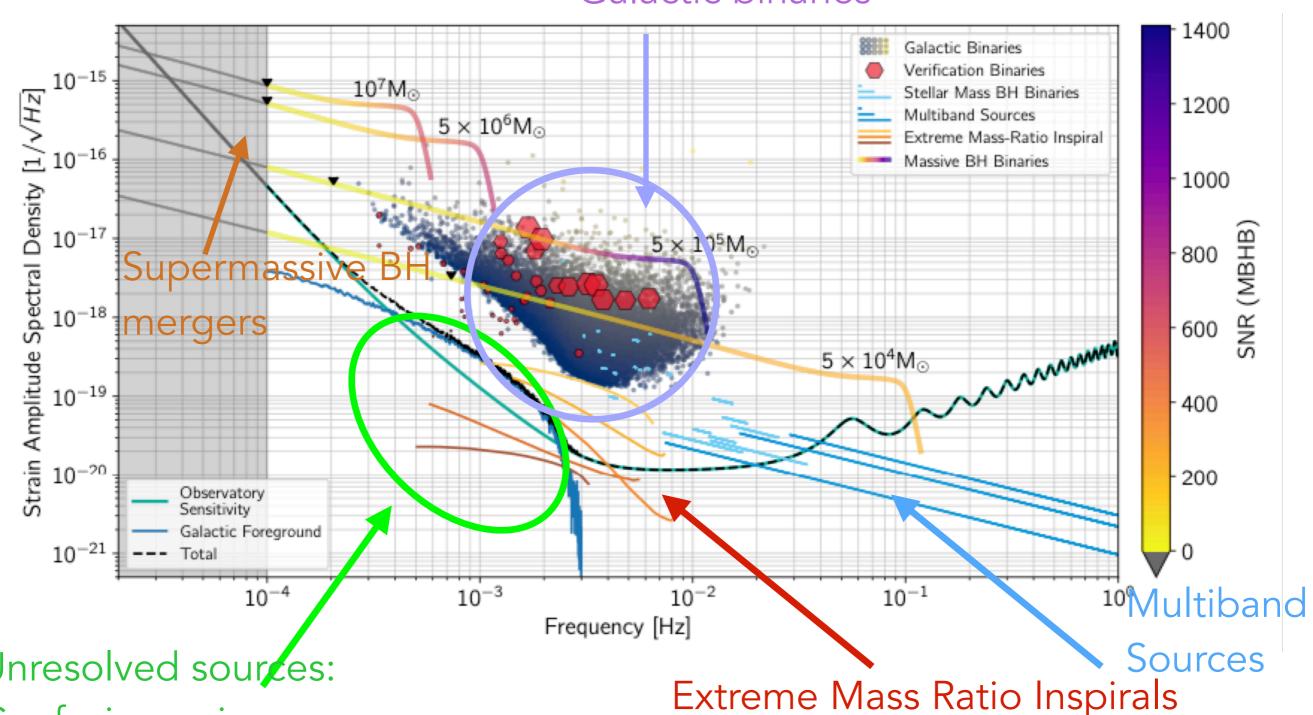


A WIDE VARIETY OF SOURCES

LISA Red Book

Transient and permanent sources

Galactic binaries



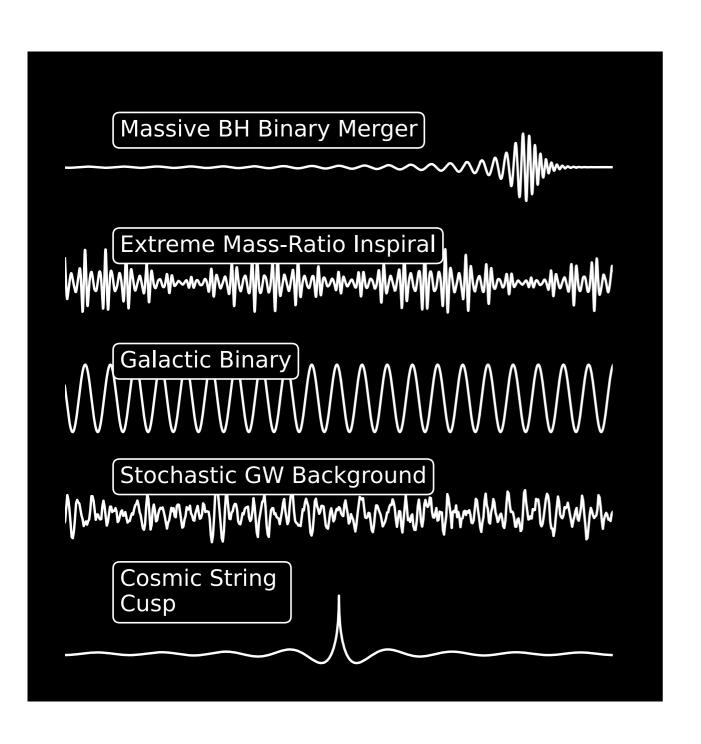
Unresolved sources:

Confusion noise





INFORMATION FROM GWS



Frequency, amplitude

Frequency evolution -> masses

Amplitude+masses -> luminosity distance

Time of arrival, amplitude, phase -> sky localisation, (very) poor in comparison to EM

Modulations of amplitude and phase -> spins & eccentricity

Effects of matter...







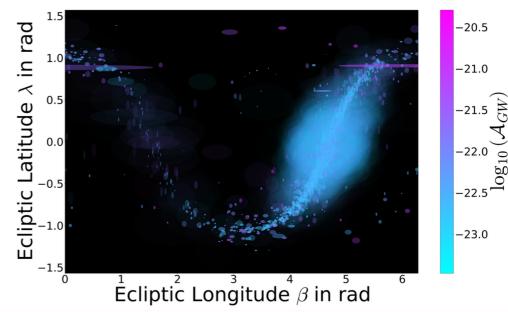


CNTS

NUMBER OF SOURCES & SKY LOCALISATION

Sources	SNR	Duration	Event rate	Sky localisation
Galactic binaries	10 – 500	permanent	10000 – 30000 detectable + background	deg^2 at best, ten deg^2 for good sources, up to 10^4 for most sources
Verification binaries	7 - 100	permanent	~40 (today)	Exact (from EM)
Stellar mass black hole binaries	7 - 30	1 to 10 years	1 to 20	Possibly < deg^2
EMRIs	7 - 60	1 year	1 to 2000 / year	< deg^2 to ~10 deg^2
Massive Black Hole binaries	10 - 3000	Hours - months	10 to 1000 / year	~ deg^2 at merger ~100 deg^4 4 hours before

Limited sky localisation -> many traditional tools not (well) adapted

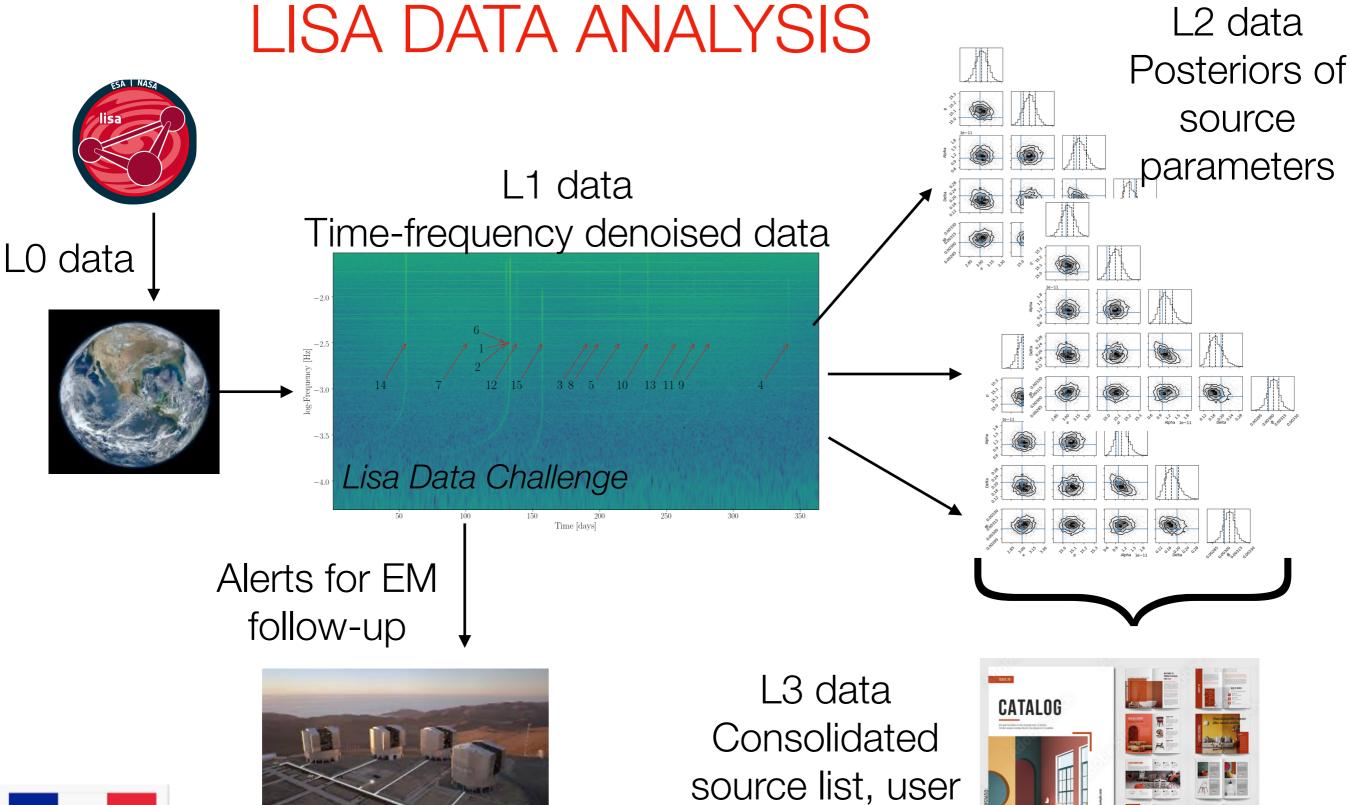














CONSORTIUM





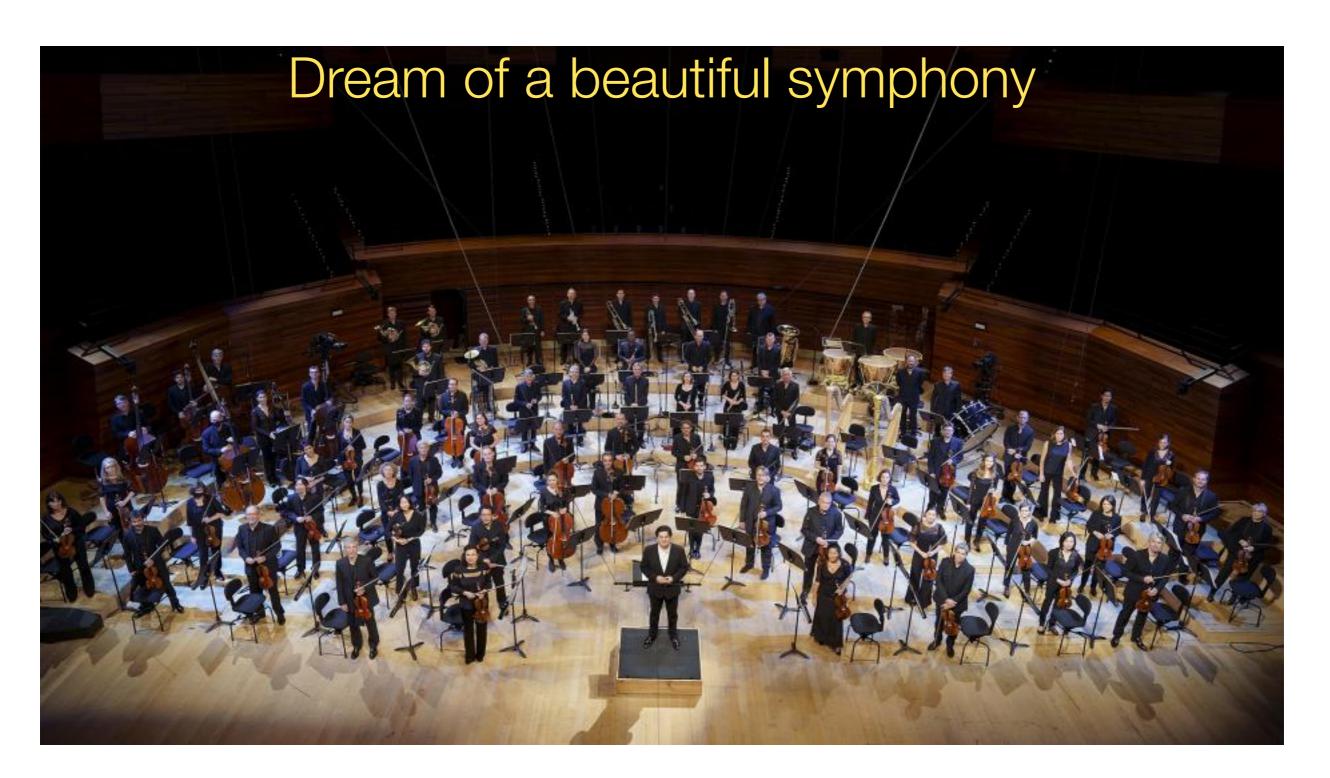






interface

LISA DATA ANALYSIS







LISA DATA ANALYSIS



Millions of superposed sources, very different from (current) ground-based detectors uncertainties are part of the pipeline -> many « ugly » uncertain PDFs



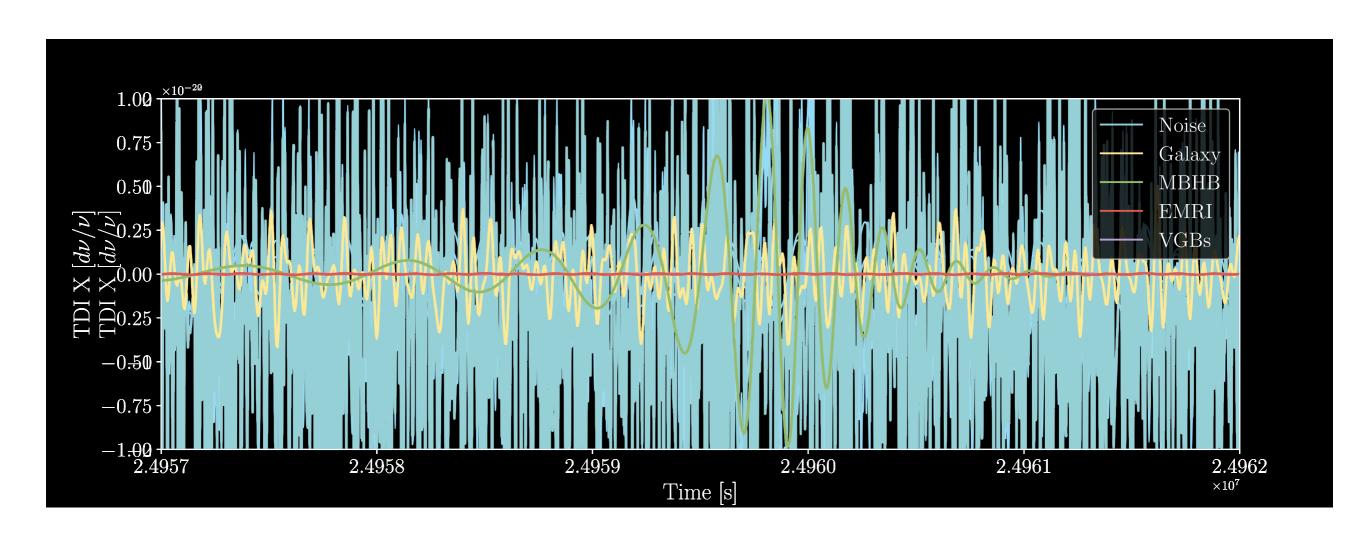






LISA DATA

LISA data challenges: time domain data



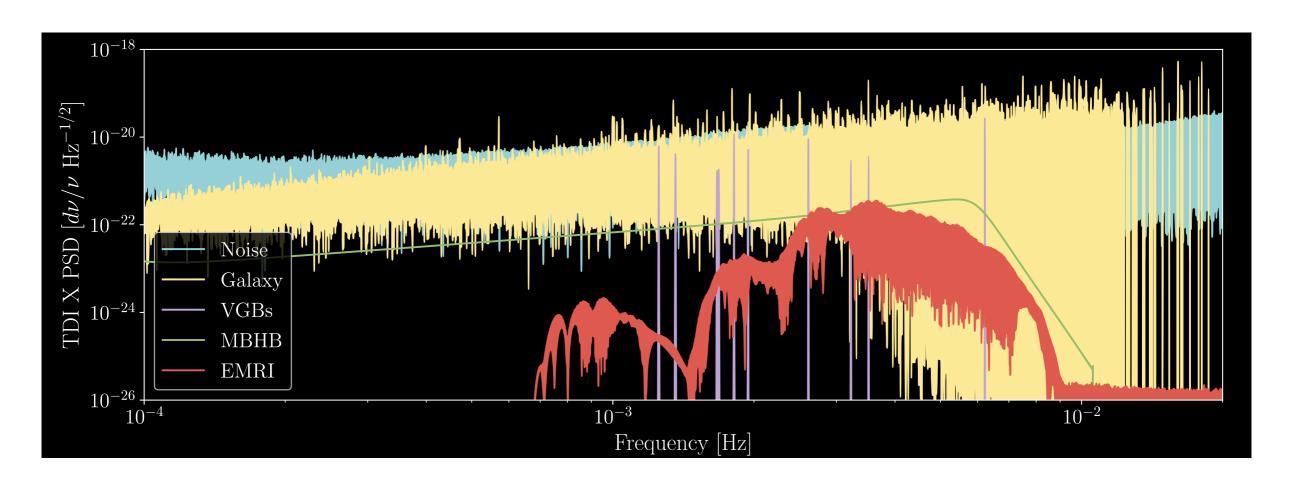
Verification GBs + EMRI + MBHB + Galaxy + Noise





LISA DATA

LISA data challenges: frequency domain data

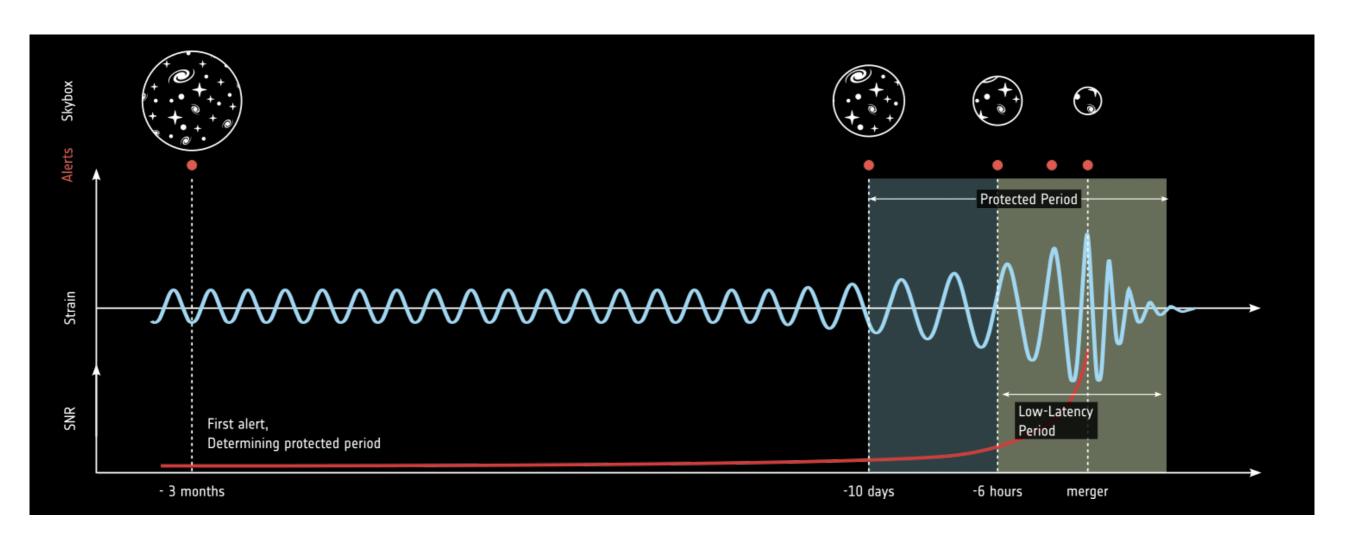


Verification GBs + EMRI + MBHB + Galaxy + Noise





ALERTS & MULTI-MESSENGER



- Several low-latency alert pipelines
 - Alerts for new sources
 - Updates on known transients

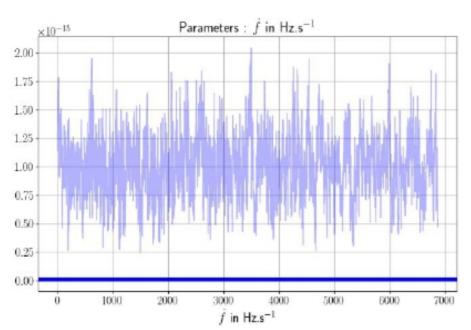




CONTENT OF DATA RELEASES (WORK IN PROGRESS)

All data levels released at the same time:

- « raw » data (combination of time series)
- denoised data: time and/or freq. domain
- several outcomes of global fits as posterior distributions, with some information on confidence
- Consolidated source list, many « uncertain » sources



 10^{16} 10^{15} f in Hz.s⁻¹

Trace of parameter in global fit -> posterior distribution







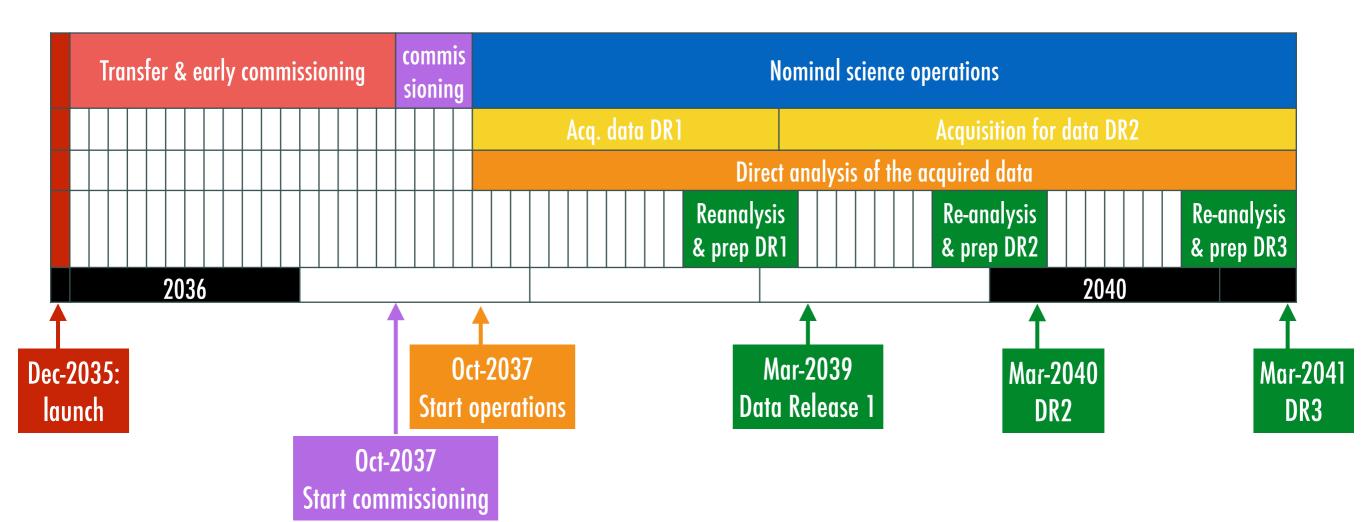


CONTENT OF DATA RELEASES (WORK IN PROGRESS)

- Noise model (very limited measurements on the ground)
- Data quality information
- Tools for X-matching with other data (but with bad sky localisation)
- => need to make it as compatible as possible with astronomy data standards, possible need for new standards



EXAMPLE OF DATA RELEASE SCENARIO



Science Management Plan:

- First DR 12 months after end commissioning
- Then, major (1/yr) and minor releases (2/yr)
- All data levels are released at the same time





KEY MILESTONES FOR DATA RELEASES

- Operations start 2037+
- First DR ~2038
- 3 mock data challenges with mock data release of increasing complexity (TBD)
 - Early 2028
 - Early 2030
 - Mid 2032
- => we want to approach/think about astro standards as soon as possible
- => looking for input and help!





